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THE CHEMISTRY OF COMETS

An Annotated Bibliography

THE CHEMISTRY OF COMETS

An Annotated Bibliography

by

J. B. Evans and V. J. DeCarlo

Melpar, Incorporated 3000 Arlington Boulevard Falls Church, Virginia

March 1965

NOTE: This work was performed on Contracts NASw-589 and NASw-890 with the National Aeronautics and Space Administration, Washington, D. C.

FOREWORD

All investigations of cometary phenomena are directed toward the elucidation of the origin, formation, and evolution of comets which will, in turn, lead to a better understanding of the origin, formation, and evolution of the Solar System. Such investigations have encompassed many disciplines and recent advances in the fields of low-temperature physics and chemistry, photochemistry, magnetohydrodynamics, charge-transfer reactions, and scattering phenomena may hasten the ultimate explanation of cometary phenomena.

The multitude of astronomical observations made in the past have allowed the analysis of the observed cometary spectra and the tabulation of the physical properties and behavior of comets. From this data, notable experimental and theoretical studies have evolved.

The experimental studies have been pointed primarily toward an explanation of the occurrence of the observed emissions. Many preliminary assignments of cometary spectra have been verified by definite identification of the emitting species through laboratory investigation. With this done, the experimental studies have been aimed toward an explanation of the chemical processes and mechanisms responsible for the formation of the emitting species.

Various theoretical studies of cometary phenomena have been advanced to explain the existence of the nucleus, the mechanisms which produce the coma, and the properties of the tail. Such theories have described the nucleus, coma, and tail in terms of structure, composition, temperature, evaporation, photodecomposition, and interaction with protons, electrons, and radiowaves. However, these theories must remain speculative as they have not, as yet, been wholly verified by experimental data.

In this respect, continued investigations in the chemistry of comets will be an invaluable tool in determining (1) the nature, concentration, ionization, and dissociation of the parent molecules of the emitting species in the coma and tail, (2) the release mechanisms responsible for the existence of the parent molecules, and (3) the composition of the nucleus.

The references included in this bibliography were accumulated at Melpar, Inc. during the course of Contracts NASW-589 and NASW-890, a program in comet chemistry directed toward the determination of the parent compounds responsible for the emissions observed in cometary spectra. Thus, all references reflect these investigations in their subject matter or represent the necessary background knowledge.

For convenience, the references are divided into three areas - (1) those which report the results of land-based observations and the theoretical interpretations resulting from these observations, (2) those which report data obtained in the laboratory as a direct result of experimentation, and (3) those which may indicate the trend of our future space investigations.

The availability of the Government contract reports is indicated when practical or necessary by the "N", "OTS", or "AD" numbers given after the reference. These numbers refer to the library indexes of the National Aeronautics and Space Administration, The Office of Technical Services, and the Defense Documentation Center, respectively.

Both an author index and a chemical index have been added to the bibliography for the user's convenience. The chemical index includes those chemical species cited in the abstracts and titles of the publications. The authors feel that, while no bibliography can claim to be complete, this compilation will provide a good basis for anyone interested in undertaking a study in the chemistry of comets.

The Chemistry of Comets -- An Annotated Bibliography

J. B. Evans and V. J. DeCarlo

CBSTRVATIONS AND INTERPRETATION

1. B. M. Middlehurst and G. P. Kuiper (Editors)
THE SOLAR SYSTIM. VOL. TV. THE MOON, METEORITES, AND COMETS
The University of Chicago Press, Chicago, Ill., 1963

A comprehensive survey of information is contained in the six chapters devoted to comets and authored by noted authorities. Chapter 15 is authored by E. Roemer and deals with the discovery, orbits, and observation of comets; Chapter 16, by J. G. Porter, with orbital statistics; Chapter 17, by K. Wurm, with the physical and chemical properties and includes discussions of the form, spectra, excitation, densities, polarization, and mechanical theory; Chapter 18, by L. Biermann and Rh. Lust, with the structure and dynamics of cometary tails; Chapter 19, by F. L. Whipple, with the structure of the nucleus; and Chapter 20, by J. H. Oort, with empirical data on the origin of comets.

2. R. A. Lyttleton
THE COMETS AND THEIR ORIGIN
Cambridge University Press, New York, N. Y., 1953

The properties and formation of comets are described in an effort to elucidate the origin of comets.

3. N. B. Richter
THE NATURE OF COMETS
Methuen and Company, Ltd., London, England, 1963

A translation of the original "Statistik und Physik der Kometen," the book provides a review of information concerning the history, statistics, theoretical models, structure, origin, formation, and decay of comets. Recent data obtained from the observations of Comets Arend-Roland and Arkos and a chapter discussing future cometary investigations are also included.

4. P. W. Merrill
SPACE CHEMISTRY
The University of Michigan Press, Ann Arbor, Michigan, 1963

The chemistry of the solar system, stars, and nebulae is discussed and includes a short treatment of the chemistry of comets.

5. P. Swings
CONSIDERATIONS REGARDING COMETARY AND INTERSTELLAR MOLECULES
Astrophysical Journal 95, 270 (1942)

A comparative study is made of the possible mechanisms of formation, dissociation, and ionization in comets and in interstellar space. The additional identification work necessary to an explanation of these processes is set forth, while the identification, ionization, dissociation, and abundance of CH^+ , CN , and C_2 are treated at length.

6. P. Swings
MOLECULAR BANDS IN COMETARY SPECTRA. IDENTIFICATIONS
Rev. Mod. Phys. 14, 190 (1942)

A discussion summarizes the identification of the molecules OH, NH, CN, CH, C_2 , CH^+ , CO^+ , and N_2 . The source of these molecules and the problems associated with their identification are treated.

7. P. Swings
COMETARY SPECTRA, REPORTS ON THE PROGRESS OF ASTRONOMY
Monthly Notices of the Royal Astronomical Society 103, 86 (1943)

A review of cometary astronomy discusses the visible, ultraviolet, astronomical, and laboratory observation of the diatomic molecules OH, NH, and CH⁺; the triatomic CH₂; and the less definite OH⁺ and NH₂. The previously identified molecules CN, C₂, CH, CO⁺, and N₂ are also treated in a concluding discussion of the excitation mechanism and rotational structure of cometary bands, and of the physical form and stability of cometary materials.

8. P. Swings and T. Page
THE SPECTRUM OF COMET BESTER (1947k)
Astrophysical Journal 111, 530 (1950)

An analysis of the ultraviolet, visible, and near infrared spectra of comet Bester (1947k) is presented with a summary of identifications for CN, C_2 , OH, NH, CO^+ , N_2^+ , OH $^+$, NH $_2$, and the λ 4050 band system. Molecules identified in the tail include CO^+ , and CO_2^+ , and the tentatively assigned NO and O_2 . Slit spectrograms of the tail cover the ultraviolet region, while spectrograms of the head cover the visible and near infrared region. A comparison of this spectra with those of other comets is given, and the transitions and origin of the specific emissions are postulated.

9. T. Page
RECENT STATISTICAL STUDIES IN ASTRONOMY
Science 132, 1870 (1960)

Statistical studies reveal information about the mass, expansion, and radio emission of galaxies and the origin and loss of comets.

10. P. Swings and L. Haser
ATLAS OF REPRESENTATIVE COMETARY SPECTRA
The University of Liege, Institute of Astrophysics, Cointe-Schlessin,
Belgium, 1956
AD 119 234

A compilation of 350 representative spectra is presented with description, notes, and references given as necessary. A foreword describes the

meaningful aspects of cometary spectroscopy and includes definitions; instrumental influences; principles for the identification of emissions; wavelength tables; the cometary bands, continuum, molecules, etc; and the interrelation between cometary spectra and that observed in the atmosphere and stars.

11. P. Swings and L. Haser
NOTES ON COMETARY SPECTRA
The University of Liege, Institute of Astrophysics, Cointe-Schlessin,
Belgium, 1956
AD 119 235

Notes published in addendum to Swing's Atlas explain the limitations effected by instrumental and observational errors and by the choice of comet model.

12. L. Haser
LA CONSERVATION DES RADICAUY LIBRES A' BASSE TEMPERATURE ET THE STRUCTURE
DES NOYAUX DE COMETES
Comptes Rendus 241, 742 (1955)

It is theorized that the observed radicals are formed by photodissociation of the parent molecules produced by sublimation of the ices in the nucleus.

13. L. Haser
SUR LA PRODUCTION DES RADICAUX LIBRES DANS LES COMETES
Bull. Acad. R. Belg. 42, 8°, 813 (1956)

A new process for the production of radicals by comets is proposed.

14. P. Swings
THE SPECTRA OF THE COMETS
Vistas in Astronomy 2, 958-981, Pergamon Press (1956)

A presentation of the general characteristics of comets is followed by a description of the ultraviolet, visible, and near infrared spectrum of the coma and tail. The assignment of the emissions is discussed along with the possible causes of excitation and the mechanisms.

15. P. Swings et al
NOTES ON COMETARY PHYSICS
The University of Liege, Institute of Astrophysics, Cointe-Schlessin,
Belgium, 1957
AD 117 169

Notes prepared preliminary to "Research on Relations between Cometary, Solar, and Atmospheric Physics," discuss the identification of emissions

in cometary spectra, the possibility of fluorescence excitation in comets, the chemical reactions in comets, and the monochromatic intensity distribution in a comet's head.

P. Swings
RESEARCH ON RELATIONS BETWEEN COMETARY, SOLAR, AND UPPER ATMOSPHERIC PHYSICS
The University of Liege, Institute of Astrophysics, Cointe-Schlessin,
Belgium, 1958
AD 152 525

An attempt to discover an interrelationship between cometary, solar, and atmospheric processes is presented in three sections. The section on cometary physics includes a number of studies: the assignment of the infrared system of CN, the identification of the forbidden lines of O(I) in comets and a comparison with that of the upper atmosphere, assignment of accurate wavelengths in the visible region and in particular for NH₂, a comparison of Comet Encke in 1957 and 1947-1937, the origin of the continuous spectrum of comets, the intensity and polarization of light scattered by cometary particles, trapped radicals and their possible reactions, and the theoretical distribution of the molecules in the head.

17. P. Swings et al
ON THE ORIGIN OF THE CONTINUOUS SPECTRUM OF COMETS
The University of Liege, Institute of Astrophysics, Cointe-Schlessin,
Belgium, 1959
AD 243 271

The possible causes of the continuous spectrum of comets are examined. Four mechanisms are considered -- diffuse reflection by the solid nucleus, scattering by molecules, scattering by solid particles, and scattering by free electrons. It is concluded that the solar continuum of comets is due primarily to scattering by solid particles.

18. P. Swings et al
RELATIONS BETWEEN COMETARY AND SCLAR ACTIVITY
The University of Liege, Institute of Astrophysics, Cointe-Schlessin,
Belgium, 1960
AD 243 274

The behavior of dust and gaseous cometary heads is compared to solar activity to indicate that the brightness of a dust head is not related appreciably to solar activity while a gaseous head may be.

19. P. Swings
EXCITATION AND ABUNDANCES IN COMETS
University of Liege, Institute of Astrophysics, Cointe-Schlessin, Belgium, 1960
AD 243 275

Cometary spectra is analyzed in studies of a fluorescence excitation

mechanism for the C₂ Swan bands, the radii of the nuclei of the periodic Comets 1957c and 1959b, and the relative abundances of CO⁺ in the tail and CN in the head of Comet Bester 1947k.

20. P. Swings, D. Bosman-Crespin and C. Arpigny
THEORETICAL CONSIDERATIONS ON THE PREDICTED FAR ULTRAVIOLET SPECTRA OF
COMETS AND OF POLAR AURORAE
University of Liege, Institute of Astrophysics, Cointe-Schlessin, Belgium,
1960
AD 257 098

The expected far ultraviolet spectra of comets and aurorae is discussed and specific predictions are tabulated for application to planned rocket and space probe investigation.

21. P. Swings
RECENT PROGRESS IN COMETARY SPECTROSCOPY
University of Liege, Department of Astrophysics, Cointe-Schlessin, Belgium, 1963
AD 430 639

The report discusses recent progress in cometary observations and spectroscopy as well as in the understanding of the excitation mechanism of the diatomic and triatomic radicals and ions.

22. A. H. Delsemme and P. Swings
HYDRATES DE GAZ DANS LES NOYAUX COMETAIRES ET LES GRAINS INTERSTELLAR
Ann. d'Astrophys. 15, 1 (1952)

The presence of solid hydrates of $\mathrm{CH_{ll}}$, $\mathrm{CO_{2}}$,... in cometary nuclei and interstellar grains is assumed and substantiated by noting that the vapor pressures of the various hydrates are of the same order of magnitude, while the vapor pressures of solid $\mathrm{CH_{ll}}$, $\mathrm{H_{2}O}$ etc. differ by very large factors.

23. P. D. Jose and P. Swings
THE SPECTRUM OF COMET 19481
Astrophysical Journal 111, 41 (1950)

From an analysis of eight spectrograms of Comet 19481, sodium emission is identified at r = 0.73 and r = 0.79 AU, and CN and the λ 4050 group at r = 2.21 AU. The usual complex structure of the CN bands is attributed to the fluorescence excitation by solar radiation. The behavior of CH and "CH₂" are compared, and unidentified features observed within and between the Swan bands and in the yellow-red region are discussed.

24. P. Swings et al
RESEARCH ON RELATIONS PETWEEN COMETARY, SOLAR AND UPPER ATMOSPHERIC PROCESSES
University of Liege, Department of Astrophysics, Cointe-Schlessin, Belgium,
March 31, 1964
AD 601 136

This report presents a discussion of the progress achieved in a program for the laboratory production of molecules of astronomical interest; study of the physical mechanisms of comets and the upper atmosphere; and investigations related to space experiments, which include far ultraviolet spectra of stars, comets and aurorae, and artificial comets.

25. L. Houziaux

DENSITY OF C₂ MOLECULES IN THE HEAD OF COMET MRKOS 1955e

The University of Liege, Institute of Astrophysics, Cointe-Schlessin,

Belgium, 1961

AD 269 914

The energy emitted in the (1,0) band of the Swan system for C_2 , as derived from photoelectric observations, is used to calculate the variation between the density of the C_2 molecule in regions close to the nucleus and that in external regions of the head.

26. P. Swings and J. L. Greenstein
PRESENCE DES RAIES INTERDITES DE L'OXYGEN DAMS LES SPECTRA COMETAIRES
Comptes Rendus 246, 511 (1958)

The identification of O(I)in cometary spectra is presented.

27. J. L. Greenstein and C. Arpigny
THE VISUAL REGION OF THE SPECTRUM OF COMET MRKOS (1957d) AT HIGH RESOLUTION
Astrophysical Journal 135, 892 (1962)

The emission line and band spectrum of the head of Comet 1957d are analyzed and presented in a table of identifications. The $\rm C_2$ and $\rm NH_2$ bands produced by resonance fluorescence were nearly completely resolved, while the $\rm Na(I)$ lines were found to be asymmetric with respect to the nucleus.

28. J. L. Greenstein
THE SPECTRUM OF COMET HUMASON (1961e)
Astrophysical Journal 136, 688 (1962)

Spectroscopic observations of Comet 1961e are analyzed and presented in a table of identifications.

29. J. L. Greenstein
HIGH-RESOLUTION SPECTRA OF COMET MRKOS (1957d)
Astrophysical Journal 128, 106 (1958)

A new second-order effect observed in high-dispersion <code>Coude'</code> spectra of <code>Comet Mrkos</code> is presented in a detailed study of <code>Swings'</code> resonance-fluorescence mechanism. Intensity ratios between individual rotational cyanogen lines were found to change within $4000 \, \text{km}$ on opposite sides of the nucleus and are reportedly caused by motions.

30. P. Swings, C. T. Elvey and H. W. Babcock THE SPECTRUM OF COMET CUNNINGHAM, 1940c Astrophysical Journal 94, 320 (1941)

The identification of well-resolved ultraviolet bands due to OH and NH in slit spectrograms of Comet Cunningham is reported along with a description of the observed CH bands, the abundances of the OH and NH molecules, the absence of CN bands in the tail, and bands observed between λ 4000 and 4130.

31. P. Swings et al
RESEARCH ON RELATIONS BETWEEN COMETARY, SOLAR AND UPPER ATMOSPHERIC PROCESSES
University of Liege, Department of Astrophysics, Cointe-Schlessin, Belgium,
March 31, 1963
N63-17540

Progress is reported for investigations in time-resolved spectroscopy in the far ultraviolet, the absorption spectra of $\rm H_2$, $\rm D_2$, $\rm HD$, $\rm C_3$, $\rm SO_2^{16}$, $\rm SO_2^{18}$, and $\rm H_2O$, and applicable efforts in instrumentation. The physical mechanisms responsible for cometary phenomena and related space investigations are discussed.

32. L. Remy-Battiau
THE LUMINOSITY VARIATIONS OF COMET HEADS
University of Liege, Department of Astrophysics, Cointe-Schlessin, Belgium,
August 20, 1964
AD 606 655

A study directed toward the elucidation of the relation between solar activity and cometary brightness re-examines Whitney's work on cometary outbursts under the assumption that the light emitted at the time of an outburst is due to scattering of solar radiation by small solid particles. The study considers the masses and required energies and concludes that solar particles are not likely to be responsible for cometary outbursts.

33. B. Donn and H. C. Urey
ON THE MCCHANISM OF COMET CUTBURSTS AND THE CHEMICAL COMPOSITION OF COMETS
Astrophysical Journal 123, 339 (1956)

Prominent comet activity is proposed to arise from explosive chemical reactions involving free radical and unstable molecules.

34. B. Donn and H. C. Urey
CHEMICAL HEATING PROCESSES IN ASTRONOMICAL OBJECTS
Mem. Soc. Roy. Sci. Liege, 4°, 18, 124 (1957)

Chemical heating by recombination of atoms, free radicals, or energetic molecules is considered as an explanation of the cause of comet outbursts.

35. B. Donn
COMETS AND THE CHEMISTRY OF MATTER IN SPACE
Astronomical Journal 64, 126 (1959)

A brief note points out that comets may be looked upon as the initial stages of accumulation from the solar nebula which could have ultimately developed into asteroidal and planetary objects and, thus, that the investigation of comets is a valid approach to cosmic chemistry.

36. B. Donn
FORMATION AND TRAPPING OF FREE RADICALS (A. M. BASS AND H. P. BROIDA, EDITORS)
Chap. 11, pp. 347-363, Academic Press, Inc., New York, N. Y., 1960

A brief description of comets is given in a review of the low-temperature chemistry of cometary constituents, comet models, cometary spectra, and postulated mechanisms or causes of cometary phenomena.

37. B. Donn
THE CHARACTERISTICS OF DISTANT COMETS
Ann. d'Astrophys. 25, 319 (1962)

The spectroscopic observations and colors of comets are analyzed and indicate that only sunlight scattered from solid grains can explain the data. The use of previous comet-tail models as set forth by Brandt, Osterbrock, Chamberlain, and Parker to explain interplanetary medium or as a solar-system probe is refuted.

38. H. C. Urey
COMETARY COLLISIONS AND TEXTITES
Nature 197, 228 (1963)

The proposed formation of tektites by cometary collision with the earth is supported.

39. B. Donn

THE ORIGIN AND NATURE OF SOLID PARTICLES IN SPACE
Presented at the New York Academy of Sciences Conference on Cosmic Dust,
Goddard Space Flight Center, Greenbelt, Md., November 21-22, 1963

An analysis of the types of particles that exist in interplanetary space is concerned with the collection and detection of such particles and with the identification of their primary source. Possible sources of such material are considered to be primary grains, cometary dust, asteroidal dust, meteorite ablation products, meteoric residue, and interstellar grains.

40. B. Donn and G. W. Sears
PLANETS AND COMETS: ROLE OF CRYSTAL GROWTH IN THEIR FORMATION
Science 140, 1208 (1963)

Crystal growth theory is applied to the formation of smoke particles from the primordial solar nebula to indicate that solid particles with filamentary structures will form and thus facilitate successive aggregation into planets, comets, and asteroids.

41. B. Donn
THE ORIGIN AND STRUCTURE OF ICY COMETARY NUCLEI
Icarus 2, 396 (1963)

The origin and structure of the nuclei is examined within the boundaries of presently accepted models. The analysis considers the distance from the sun, velocity, temperature, cloud of comets, and the formation of planets. The structure of the nuclei is considered in terms of an $\rm H_2O$ ice, density, pressure, and temperature.

42. L. S. Marochnik
THE FORM OF COMETARY ENVELOPES
Sov. Astron. 4, 480 (1960) - Astron Zh. 37, 508 (1960)

Alfven's hypothesis relating to the interaction between cometary gas and a stream of particles carrying a magnetic field is used to show that the form of a cometary envelope is described by a catenary, in accord with observations. The parabolic relationship flowing from the mechanical theory of cometary forms is only a rough approximation. The transverse motions observed in the tails of comets are accounted for in terms of Alfven waves. The upper limit of the magnetic field is estimated. (Author)

43. L. S. Marochnik
SOME MECHANISMS OF THE RADIO EMISSION OF COMETS
Byul. Komis. po. Kometam i Meteoram Astron. Soveta AN SSSR, No. 5, 21-27
(1961): Ref. Zh., Fizika, No. 1, 49 (1963), Abstract 1Zh294
AD 412 130

An investigation is carried out of the transformation of cometary gas plasma oscillations into radio emission. The effective temperature of the plasma waves and the intensity of the arising radio emission are calculated. The latter corresponds to the observed values. It is shown that the reflection of the solar radio emission by the comet plasma is inessential. The effect of the comet's proper magnetic field is also inessential. (LM)

44. L. S. Marochnik ON THE PLASMA NATURE OF A COMETS' HEAD Sov. Astron. 6, 532 (1963) - Astron. Zh. 39, 678 (1962)

The gas in a comet's head is considered to be a plasma with a high degree of ionization. The shape of the comet's head and the origin of the ionization are also treated.

L. S. Marochnik
INTTRACTION OF SCLAR CORPUSCULAR STREAMS WITH COMETARY ATMOSPHERES.
I. SHOCK WAVES IN COMETS
Sov. Astron. 6, 828 (1963) - Astron. Zh. 39, 1067 (1962)

An attempt is made to develop the theory of shock waves in application to comets. The possibility of considering the process of collision of a cometary head with a solar corpuscular stream in terms of magnetohydrodynamics is shown. The parameters of shock waves originating during this process are computed. The results obtained are used to analyze some phenomena in comets. (Author)

46. L. S. Marochnik
WAVE MOTIONS IN COMET TAILS
Sov. Astron. 7, 218 (1963) - Astron. Zh. 40, 284 (1963)

The causes for wave motions in ionized cometary tails are discussed. It is shown that these motions are related to magnetohydrodynamic waves. The increase in amplitude and wavelength with distance from the head to the tail is accounted for as an increment in the local Alfven velocity.

47. L. S. Marochnik
INTERACTION BETWEEN SOLAR CORPUSCULAR STREAMS AND COMETARY ATMOSPHERES.
II. "COLLAPSING" ENVELOPES. RADIO-FREQUENCY EMISSION
Sov. Astron. 7, 334 (1963) - Astron. Zh. 40, 504 (1963)

The collapsing envelopes observed in Comet Morehouse 1908 III and other

comets are analyzed and correlated with the occurrence of ionization and radio-frequency emission.

48. L. S. Marochnik CAUSE OF IONIZATION OF COMETARY MOLECULES Sov. Astron. 7, 544 (1964) - Astron. Zh. 40, 714 (1963)

In this paper it is shown that the focussing of the magnetic field of the solar corpuscular stream leads to an increase in the concentration of protons in the stream. Even if the corpuscular stream was extremely rarified before collision with the comet, the increase of proton concentration is sufficient for explaining the observed lifetime of parent molecules before ionization by recharging $H^+ + M - M^+ + H + M^- + M^+ + H + M^- +$

49. V. I. Cherednichenko THE DISSOCIATION AND IONIZATION OF COMETARY MOLECULES IN THE PHOTON AND CORPUSCULAR FIELD OF SOLAR RADIATION Astron. Zh. 36, 254 (1959)

An attempt is made to determine the lifetimes of prechosen cometary molecules under bombardment by photon and corpuscular solar radiation.

50. V. I. Cherednichenko LIFETIME OF COMETARY ICES IN THE FIELD OF PHOTON AND CORPUSCULAR SOLAR RADIATION Meteoritika, Akad. Nauk. SSSR, Komitet Meteorit. No. 19, 143-54 (1960)

A more complete description is given for the calculation of the lifetimes of assumed cometary molecules under the influence of photon and corpuscular radiation.

51. Yu. V. Evdokimov MOTION OF THE JACOBINI-ZINNER COMET FROM 1933 TO 1946 Sov. Astron. 7, 415 (1963) - Astron. Zh. 40, 544 (1963)

The orbital elements of the Jacobini-Zinner comet are computed from observations made of the 1933 and 1939 apparitions and the 1939 and 1946 apparitions.

52. Y. V. Yevdokimov

MASSES OF COMET GIACOBINI-ZINNER AND THE DRACONID METEOR STREAM Smithsonian Contributions to Astrophysics, Vol. 7, pp. 297-299, Smithsonian Institution Astrophysical Observatory, Washington, D. C., 1963 N63-16664

The change in the mass and motion of Comet Giacobini-Zinner is assessed and related to the Draconid meteor stream.

53. V. Fessenkov

A NOTE ON THE COMETARY NATURE OF THE TUNGUS METEORITES Smithsonian Contributions to Astrophysics, Vol. 7, pp. 305-308, Smithsonian Institution Astrophysical Observatory, Washington, D. C., 1963 N63-16664

The fall of the Tungus meteorite is analyzed with evidence given to support the theory that it was a small comet.

54. S. K. Vsekhsvyatskiy

THE PROBLEMS ABOUT COMETS AND THE SOLAR SYSTEM
Transactions of the Congress of the All-Union Astronomic-Geodetic Society,
Moscow, 1960, pp. 107-119 (1962)
OTS/SLA 63-19537

The theories for the origin of comets and the mechanism of their formation are reviewed. It is concluded that comets and other small bodies permit an estimation of the amount of matter ejected from planets and thus, the study of comets will allow the history of the planetary system to be elucidated.

55. S. K. Vsekhsvyatski

ABSOLUTE MAGNITUDES OF 1954-1960 COMETS Sov. Astron. 6, 849 (1963) - Astron. Zh. 39, 1094 (1962)

The absolute magnitudes and photometric parameters of 62 comets observed during 1954-1960 are compiled from available photometric and brightness data.

56. S. K. Vsekhsvyatskii

OBSERVATIONS AT KIEV OF COMET SEKI-LINES 1962 Sov. Astron. 7, 133 (1963) - Astron. Zh. 40, 176 (1963)

The brightness of the head of Comet Seki-Lines 1962 covering the entire apparition from April 3 to May 6, 1962 is presented.

57. B. Yu. Levin
THE STRUCTURE OF ICY COMET NUCLEI
Sov. Astron. 6, 593 (1963) - Astron. Zh. 39, 763 (1962)

A hypothesis is proposed that the substances of the icy cometary nuclei are present in the form of separate atoms and molecules embedded in an amorphous noncoherent condensate of different volatile substances, and secondly, that the evaporation of volatile substances from the nucleus leaves a porous matrix, the fragments of which are meteoric particles.

58. V. M. Yuzhakov REMARKS ON THE MAGNETIC FIELD IN THE TAIL OF THE COMET 1957d Sov. Astron. 7, 591 (1964) - Astron. Zh. 40, 779 (1963)

The induction of the magnetic field in the tail of the Comet Mrkos 1957d is determined from the helical structure of the field, and the value found was 5.10-6 G, which is less than the interplanetary field by a factor of approximately ten. (Author)

59. D. O. Mokhnach

EFFECT OF PHOTODISSOCIATION ON THE DISTRIBUTION OF SURFACE BRIGHTNESS IN

THE CARBON ATMOSPHERE OF COMETS

Dokl. AN SSSR, Astronomiya, Tom 157, No. 2, 309-312, Izdatel'stvo "Nauka",

1964

ST-OA-10195 (NASA)

An attempt is made to construct a simple model of a carbon coma, considering the deviation in the distribution of surface brightness (1/r and $1/r^2$) as observed in Comet Burnham (1959k).

60. K. A. Shteins and S. Ya. Sture
THE DIFFUSION OF COMETS
Sov. Astron. 6, 398 (1962) - Astron. Zh. 39, 506 (1962)

As Part IV, the problem of new comets is assessed in terms of diffusion processes. The study concludes that the stabilization of the diffusion process involves several million years.

61. E. A. Dibai
ORIGIN OF COMETARY NEBULAE II
Sov. Astron. 7, 606 (1964) - Astron. Zh. 40, 795 (1963)

It is shown that the principal morphological characteristics of cometary nebulae may be represented as the result of focusing of shock waves originating at the interface between a cold, dense neutral gas and an ionized gas in the Stromgren zone. (Author)

62. O. V. Dobrovol'skiy
RADIO EMISSION FROM COMETS
Byul. in-ta Astrofiz. AN Tadzh SSR, No. 26, pp. 3-11 (1958)
AD 264 497

A review of the reported radio emissions observed from comets and the theoretical mechanism or cause of such emission is given.

63. A. D. Dubyago STRUCTURE OF COMET NUCLEI AND FORMATION OF METEOR STREAMS Astron. Zh. 27, 5 (1950) AD 120 694

The structure of comet nuclei is considered as consisting of many separate solid bodies of a quite large size. It is shown that the density of the nuclei can be estimated from the periodic processes observed in comets. The velocity with which meteors leave the comet head is estimated. It is concluded that the meteor swarms associated with periodic comets must originally have a small extent perpendicular to the plane of the orbit, and that the rate of propagation of the swarm may be evaluated.

64. V. G. Fesenkov
ON THE NATURE AND ORIGIN OF COMETS
Sov. Astron. 6, 459 (1963) - Astron. Zh. 39, 583 (1962)

In an attempt to discover the origin of comets, it is postulated that short-period comets are generated from aperiodic comets, that cometary nuclei consist of compact agglomerations of fairly unstable particles, that the total mass of a comet can be estimated from the brightness of the comets head and from reaction effects attendant upon the rotation of the head about the comets axis, and that the distance between the component parts of the nucleus is quite short. In conclusion, it is stated that such compact clusters of unstable species could not possibly originate in volcanic eruptions from the interior of a planet.

65. B. A. Vorontsov-Vel'yaminov
PARTIAL GAS DENSITIES IN COMET 1942g WHIPPLE
Sov. Astron. p. 674 - Astron. Zh. 37, 709 (1960)

The partial gas densities in Comet Whipple are computed from the absolute monochromatic emission spectra of the CN 3883, $\rm C_2$ 4730, and $\rm C_3$ 4050 bands as determined by reference to stellar spectra.

66. S. M. Poloskov
THE INTRARED SPECTRUM OF COMETS
ca 1951
N64-23179

A study of the infrared spectra of comets refutes Swings identification of the (red) system of the CN molecule and assigns this emission to $A^2\pi-2\Sigma$ system of the N₂⁺ molecule.

67. S. V. Orlov and S. M. Poloskov COMETS
?
N64-24654

A summary of Soviet studies of comets covers comet models, cometary physics, and the origin and evolution of comets.

68. L. Biermann
COMET TAILS AND SOLAR CORPUSCULAR RAYS
Zeitschrift fur Astrophysik 28-29, 274 (1950-51)

The acceleration of long comet tails, consisting of CO⁺ and other ions, to that which exceeds the gravitational acceleration of the sun is theorized to be produced by the solar corpuscular radiation of the sun.

69. L. Bierman and E. Trefftz

UBER DIE MECHANISMEN DER IONISATION

Zeitschrift fur Astrophysik 59, 1 (1964)

The possible mechanisms for the excitation of the forbidden oxygen lines in comets are discussed along with the production by ionization of suitable parent molecules which can account for the observed emissions of C_2 , CN, and CO^+ .

70. D. Antrack, L. Bierman and Rh. Lust SOME STATISTICAL PROPERTIES OF COMETS WITH PLASMA TAILS Annual Review of Astronomy and Astrophysics, Vol. 2, p. 327, Annual Reviews, Inc., Palo Alto, California, 1964

The report surveys the results obtained from the observations of comets in a study of the solar plasma through its effects on the plasma-tail comets.

71. N. T. Bobrovnikoff
PHYSICAL THEORY OF COMETS IN THE LIGHT OF SPECTROSCOPIC DATA
Rev. Mod. Phys. 14, 164 (1942)

A brief summary of the general problems raised by new spectral data is presented in order to point out certain unacceptable physical theories and to outline observational procedure which may serve to provide better understanding of cometary phenomena.

72. N. T. Bobrovnikoff
ON THE SPECTRA OF COMETS
Astrophysical Journal 66, 439 (1927)

The spectra of 22 comets obtained between 1908 and 1927 are analyzed. A correlation is drawn between the heliocentric distance and the type of continuous spectra and between the phase-angle and the type of spectra. Sudden changes in the spectra are discussed and reasons in support of a fluorescence origin of cometary spectra are given.

73. J. A. Bruwer
PHOTOGRAPHIC OBSERVATIONS OF COMETS
Circulars 7, 20 (1963)

The photographic observation of Comets Seki-Lines 1962c, Temple 2, and Humason 1961e is tabulated, listing date of observation, magnitude, and position.

74. A. J. Cohen
ASTEROID- OR COMET-IMPACT HYPOTHESIS OF TEXTITE ORIGIN: THE MCLDAVITE
STREWN-FIELDS
Textites, pp. 189-212, The University of Chicago Press, Chicago, Ill.,
1963
N64-14934

Arguments are presented in support of the comet-impact terrestrial origin of tektites.

75. R. Coutrez, J. Hunaerts and A. Koeckelenbergh RADIO EMISSION FROM COMET 1956h ON 600 MC Proc. IRE 46, 274 (1958)

An attempt is made to identify the emission as due to a Λ -type doubling of the level J = 3/2 for the $^2\pi_{3/2}$ state of the CH molecule.

76. J. Dufay and A. Baranne LE SPECTRE DE LA COMETE WILSON-HUBBARD (1961d) Ann d' Astrophys. 25, 301 (1962)

The spectra of Na, NH $_2$, C $_2$, and CN as recorded photographically are discussed and explained.

77. W. C. Erickson and P. Brissenden
A SEARCH FOR DECAMETRIC RADIATION FROM THE COMET WILSON 1961d
Astrophysical Journal 136, 1138 (1962)

A search for decametric radio emission from Comet Wilson made by Clark Lake radio astronomy station reports definite negative results.

78. F. H. Flynn
INTENSE FAR-RED EMISSION FROM COMET CANDY (1960n)
Observatory 81, 247 (1961)

The observation of red emission near 7900 and 3100\AA is tentatively assigned to CN.

79. C. C. Kiess
RECENT ADVANCES IN ASTRO-GEOPHYSICS
Georgetown College Observatory, Georgetown University, Washington, D. C.,
1964
N6h-15286

A brief lecture presents a general description of comets and the role they play in the study of the formation of the solar system.

80. W. Liller
THE NATURE OF THE GRAINS IN THE TAILS OF COMETS 1956h AND 1957d
Astrophysical Journal 132, 867 (1960)

Continuous spectra between λ 3400 and λ 6400 obtained from observations of the tails of Comets Arend-Roland (1956h) and Mrkos (1957d) are compared with theoretical light scattering curves. Results indicate that iron spherules of an average diameter of 0.6 μ and masses of 8 x 10⁻¹³g produced the radiation. Considerations of the brightnesses of the tails allow estimates to be made of the total masses of the particles.

81. W. Liller
PHOTOELECTRIC PHOTOMETRY OF COMETS
Astronomical Journal 66, 372 (1961)

A description is given of narrow-band photoelectric photometry of Comets

Arend-Roland (1956h), Mrkos (1957d), Encke, and Burnham (1959k). The observations are summarized in three parts dealing with head spectra, tail spectra, and monochromatic isophotal contours of Comet Burnham.

82. R. A. Lyttleton
A COMETARY MECHANISM FOR THE FORMATION OF TEXTITES
Proc. Roy. Soc. A 272, 467 (1963)

The formation of tektites is theorized to be caused by short-period comets which have an eccentricity small enough to allow the accretion of cometary material to occur during the passage of the earth through the comet and thus, produce a narrow jet of material falling vertically downward through the atmosphere.

83. R. A. Lyttleton and J. M. Hammersley
THE LOSS OF LONG-PERIOD COMETS FROM THE SOLAR SYSTEM
Monthly Notices of the Royal Astronomical Society 127, 257 (1964)

The loss of long-period comets from the solar system by planetary action is estimated statistically.

84. B. G. Marsden and G. van Biesbroeck
THE ORBIT OF COMET 1944I (van GENT-PELTIER-DAIMACA)
Astronomical Journal 68, 235 (1963)

A hyperbolic orbit of the Comet 1944I, first observed in 1943 by Gent, is determined from positions covering a 57-day arc.

85. G. van Biesbroeck and B. G. Marsden ORBIT OF COMET 1954V (ABELL) Astronomical Journal 68, 212 (1963)

From 34 observations made a year after the comet had attained its perihelion, it is calculated that the hyperbolic excess of the osculating orbit is 19 times its mean error, but that the "original" orbit was elliptical.

86. A. McKellar

SCME TOPICS IN MOLECULAR ASTRONOMY

J. Roy. Astron. Soc. Canada 54, 97 (1960)

Spectroscopic studies of molecules in various astronomical bodies are reviewed.

87. A. McKellar

INTENSITY MEASUREMENTS ON EMISSION BANDS IN COMETARY SPECTRA Rev. Mod. Phys. 14, 179 (1942)

Intensity measurements on the bands in the spectra of Comet Jurlof-Achmarof-Hassel (1939d) and Comet Cunningham (1940c) are presented in one of the first reports on the quantitative spectrophotometry of comets. The mechanism of the excitation of the CN, C_2 , and CH cometary bands is considered in the light of such measurements. The resonance-fluorescence mechanism apparently accounts for the production of CN emission bands.

88. A. McKellar

COMPARISON OF THE 3883 CN BAND IN THE SPECTRA OF COMETS 1940c AND 1942g Astrophysical Journal 100, 69 (1944)

Differences in the structure of the λ 3883 (0,0) CN band as observed in the spectra of Comet Cunningham (1940c) and Comet Whipple-Fedtke-Tevzadze (1942g) are noted and discussed in terms of the resonance production of the bands and the effects of the different heliocentric distances and radical velocities of the comets.

89. A. McKellar

ROTATIONAL DISTRIBUTION OF CH MOLECULES IN THE NUCLEUS OF COMET CUNMINGHAM (1940c)
Astrophysical Journal 98, 1 (1943)

From the observed profile of the λ 4315 $^2\Delta$, $^2\pi$ CH band in the comet, the distribution of the rotational levels of the upper $^2\Delta$ state is found. Assuming the resonance mechanism for the production of the band, the distribution among the rotational levels of the normal $^2\pi$ state is derived. The results indicate that all the molecules exist in the two lowest rotational levels, K = 1 and K = 2.

90. F. D. Miller

FILTERS FOR COMET PHOTOGRAPHY
Pub. Astron. Soc. Pac. 70, 279 (1958)

The analysis of several components of the structure of a comet by means of direct photography with suitable emulsion/filter combinations is illustrated with results obtained from the observation of Comet Mrkos 1957d.

91. F. D. Miller

THE TYPE I TAIL OF COMET 1955e
Pub. Astron. Soc. Pac. 74, 60 (1962)

Extended analysis of observations made of Comet Mrkos (1955e) presents two previously unobserved bands of CO⁺ at λ 6139 and 6239 and an unidentified emission near λ 6600. The existence of the comet's well-developed Type-I tail observed at a high heliographic latitude in a period of comparative solar calm is brought forth as noteworthy.

92. F. D. Miller
CBJECTIVE-PRISM SPECTROGRAMS OF COMMIT HUMASON (1961e)
Pub. Astron. Soc. Pac. 74, 528 (1962)

Analysis of four objective-prism spectrograms of Comet Humason (1961e) is given. Bands of neutral molecules such as CN and C_2 were not found in the spectra. The (1,0), (2,0), and (3,0) CO^+ tail bands appeared superimposed on a continuum.

93. F. D. Miller
NOTE ON THE SPECTRUM OF COMET IKEYA (1963a)
Astrophysical Journal 139, 766 (1964)

Unidentified emission at λ 6158, 6199, 6542, and 6562 observed in the spectra of Comet Ikeya (1963a) is reported and postulated to be due to ions.

94. M. F. Walker
OBSERVATIONS OF COMETS BAKHAREV-MacFARLAND-KRIENKE, 1955f, AND BAADE, 1954h
Pub. Astron. Soc. Pac. 70, 191 (1958)

The results of three-color photoelectric observations obtained in 1955 at the Mount Wilson and Palomar Observatories are reported.

95. F. L. Whipple
A COMET MODEL. I. THE ACCELERATION OF COMET ENCKE
Astrophysical Journal 111, 375 (1950)

A comet model is proposed, in which the nucleus is visualized as a conglomerate of ices, "volatiles," and meteoric materials combined at extremely low temperatures. The model resolves the chief problems of cometary motions and accounts for a number of other cometary phenomena.

96. F. L. Whipple
A COMET MODEL. II. PHYSICAL RELATIONS FOR COMETS AND METEORS
Astrophysical Journal 113, 464 (1951)

The loss of gaseous and meteoric materials from the nucleus is considered in terms of the comet's structure and compared with previous observations.

97. F. L. Whipple
A COMET MODEL. III. THE ZODIACAL LIGHT
Astrophysical Journal 121, 750 (1955)

An attempt is made to explain the occurrence of the zodiacal light, which

arises from the scattering and diffraction of sunlight by small particles, by the postulation of a comet model which can account for the replenishment of these particles by ejection or disintegration by comets.

98. F. L. Whipple PROBLEMS OF THE COMETARY NUCLEUS Astronomical Journal 66, 375 (1961)

The icy model is examined critically through a presentation of the relevant observational and theoretical evidence of its existence and nature.

99. F. Whipple REMARKS ON COMETS, METEORS, AND PLANETARY EVOLUTION Smithsonian Astrophysical Observatory, Harvard College Observatory, Cambridge, Mass. N64-24771

A discussion of comet accumulation in the outer parts of the solar system is presented.

100. C. Whitney COMET OUTBURSTS Astrophysical Journal 122, 190 (1955)

Several comet outbursts accompanied by the ejection of spherical halos are investigated. Special attention is paid to the mass and energy ejected in a typical halo. On the basis of the results, it is suggested that the energy of an outburst may be derived from normal insolation. The activity can possibly be explained in terms of the icy-conglomerate comet nucleus. (Author)

101. A. A. Wyller C₂ PHILLIPS BANDS IN EMISSION FROM COMET CANDY (1960n)? Observatory 82, 73-5 (1962)

It is suggested that the λ 7906 and λ 8106 emission observed from Comet Candy may be due to the infrared C₂ Phillips system ($^1\pi$ u = $^1\Sigma$ g $^+$).

102. H. Zanstra THE EXCITATION OF LINE AND BAND SPECTRA IN COMETS BY SUNLIGHT Monthly Notices of the Royal Astronomical Society 89, 178 (1928)

The line and band spectra observed in the head of a comet are theorized to be produced by the absorption of sunlight and subsequent resonance or fluorescence re-emission.

103. COMETARY TAILS
Defense Documentation Center, Cameron Station, Alexandria, Va., July 1962
ARB No. 10857

A bibliography of 29 reports dealing with observational and laboratory investigations of comets is offered.

104. H. Alfven
ON THE THEORY OF COMET TAILS
Tellus IX, 92 (1957)

It is shown that some of the difficulties of Biermann's theory, which postulates that the repulsive force in comet tails is due to a corpuscular radiation from the sun, can be overcome if the assumed radiation consists of beams with a frozen-in magnetic field of the same type as required in the electric field theory of magnetic storms and aurorae. The interaction of such a beam with the head of a comet could produce an amplified magnetic field which determines the shape of the tail. The high accelerations observed in tails may be due to electromagnetic forces.

105. D. B. Beard and M. P. Nakada INTERPLANETARY MAGNETIC FIELDS AS A CAUSE OF COMET TAILS Nature 199, 580 (1963)

In explanation of the formation of cometary tails, it is pointed out that the gases in cometary comae can be efficiently ionized by a solar wind of ionized hydrogen embedded in a magnetic field, and that an interplanetary magnetic field can couple the cometary gas to the solar wind.

106. T. Carrington
FLUCRESCENCE IN COMETS AS A MARKOV PROCESS
Astrophysical Journal 135, 883 (1962)

The fluorescence mechanism for the emission of electronic spectra of diatomic radicals in comet heads, as proposed by Swings and Hunaerts, is treated statistically and compared with previous less-rigorous calculations.

107. F. Hoyle and M. Harwit
PLASMA DYNAMICS IN COMETS. I. PLASMA INSTABILITY
Astrophysical Journal 135, 867 (1962)

An attempt is made to find a mechanism consistent with Bierman's theory that tail ions are accelerated through interaction with streams of solar particles. The field-free case is examined in an effort to establish conditions for which cometary plasma acceleration may be enhanced by plasma instability. From this it is concluded that plasma instability does not appreciably contribute to tail plasma acceleration.

108. M. Harwit and F. Hoyle
PLASMA DYNAMICS IN COMETS. II. INFLUENCE OF MAGNETIC FIELDS
Astrophysical Journal 135, 875 (1962)

In an attempt to find a mechanism consistent with Bierman's theory, it is shown that, if predominantly transverse magnetic fields are imbedded in the solar stream, the observed accelerations can be readily imparted to tail ions. The 10 km sec⁻¹ ejection velocities from comet heads and the filamentary structure of many comet features are accounted for. Observational tests for the transverse field hypothesis are suggested.

109. C. R. O'Dell
EMISSION-BAND AND CONTINUUM PHOTOMETRY OF COMFT BURNHAM, 1959k
Pub. Astron. Soc. Pac. 73, 35 (1961)

Two interference filters were used to observe the continuum around λ 4470 and the C₂ Swan band sequence near λ 4700. The data obtained includes information on the variation of surface brightness with distance from the nucleus, on changes in the comet with time and heliocentric distance, and the distribution of the material responsible for the band or continuum emission.

110. D. E. Osterbrock
A STUDY OF TWO COMET TAILS
Astrophysical Journal 128, 95 (1958)

Photographic observations of the directions of the tails of Comet Baade (1954h) and Comet Haro-Chaviro (1954k) indicate that the tail lies in the orbital plane roughly midway between the radial and tangential directions. Analysis of the observations indicate that the material in the tail must be subjected to a resisting force roughly equal to the radial repulsive force of the sun, that this resisting force can be caused by the interplanetary gas if the tail contains a large fraction of hydride molecules, and that the possibility the tail consists of solid particles small in comparison with the wavelength of light can not be ruled out.

111. C. R. O'Dell and D. E. Osterbrock
EMISSION-BAND AND CONTINUUM PHOTOMETRY OF COMET SEKI (1961f)
Astrophysical Journal 136, 559 (1962)

 C_2 densities are obtained from photoelectric measurements of the continuum and C_2 emission-band fluxes from Comet Seki (1961f) and Comet Burnham (1959k).

112. J. A. O'Keefe and B. N. Shute ORIGIN OF TEKTITES Science 139, 1288 (1963)

It is shown from the observed distribution of tektites that they can not be the result of cometary or meteoritic impact but, rather, can be accounted for as a result of fusion stripping of a satellite.

113. E. J. Öpik
PHOTOMETRY, DIMENSIONS, AND ABLATION RATE OF COMETS
The Irish Astronomical Journal 6, 93 (1963)

Considering the peculiarities of light distribution in the coma, it is estimated that the brightness of comets varies inversely as the first power of distance from the observer instead of the traditionally-used second power. This is postulated to be of major importance for artificial comets, making their construction prohibitive. The sudden appearance of radicals and ions in cometary envelopes is said to point to an active shell or radiation belt surrounding the nucleus. Formulae for absolute magnitudes and photometric radii of comets are given and confirmed by recoil radii. These are seen to be much smaller than the dimensions of "false nuclei" observed telescopically. The total evaporation rate of the nuclear ices is estimated, and the luminescent radicals (C_2 , CN) and ions (CO^+) are reported to account for almost one percent of the total gaseous mass. The layered mechanism of simultaneous evaporation of constituents of different volatility is discussed. From molecular weight and cosmogonic considerations, it is theorized that solid hydrogen is an important constituent of the nuclei. The mechanism of outbursts is said to be a sudden increase of the evaporation area caused by shake-off of clouds or fragments. The revision of the photometric parameters of comets, based on new formulae, is suggested.

114. E. J. Öpik
SURVIVAL OF COMET MUCLEI AND THE ASTEROIDS
University of Maryland Mimeo, 59 pp., 1961

The article considers the dynamic survival of comets and other stray bodies in encounters with planets in an effort to uncover by statistics any possible genetic link between comets and asteroids.

115. E. J. Öpik

NOTE ON STELLAR PERTURBATIONS OF NEARLY PARABOLIC ORBITS

Proc. Amer. Acad. Arts and Sciences 67, 169 (1932)

The effect of passing stars on the stability of the orbits of meteors and comets in the solar system is assessed in a statistical consideration.

116. E. J. Öpik
THE SPIKE OF COMET AREND-ROLAND 1956h
The Irish Astronomical Journal 5, 37-50 (1958)

The spike of Comet Arend-Roland is discussed and interpreted to be a "synchrome" consisting of dust particles of different radii released in an outburst which left the nucleus instantaneously devoid of all volatile material.

117. E. J. Öpik
THE MOTION OF THE CONDENSATION IN THE TAIL OF HALLEY'S COMET JUNE 5-8, 1910
Zeitschrift fur Astrophysik 58, 192 (1964)

The observed motions of the condensation in Halley's comet are found to agree with that of hyopthetical rays calculated from the observed behavior of rays in Type-I plasma tails. It is suggested that this analogy may offer a clue to the understanding of the evolution and metamorphosis of the observed ray emission in comet tails.

118. D. H. Robey
A NEW MODEL FOR COMETS--THE COLD, PARTIALLY CONDENSED, MAGNETIZED PLASMA
J. Astronautical Sciences 9, 41 (1962)

A study of the general properties of comets presents those characteristics which suggest that a comet's nucleus may be surrounded by and attached to a magnetized plasma. This theory is used to postulate a formula for coma contraction and is also applied to comet outbursts. The contraction is believed to be caused by the solar wind which may, during periods of solar flare, cause a sudden contraction to effect a comet outburst.

119. G. J. Szasz and S. F. Singer
SYMPOSIUM ON THE PHYSICS OF COMETS AND METEORS
Office of Naval Research, London, November 18, 1952
AD 382

An account of a recent Liege symposium includes discussions of the spectroscopy of comets, the structure of the head and tail, and the origin and evolution of comets.

120. K. Wurm
THE ULTRAVIOLET SOLAR SPECTRUM AND COMETS
Proceedings of the Tenth International Astrophysical Symposium, Liege,
July 11-14, 1960
AD 269 781

The importance of the interaction of solar ultraviolet radiation with cometary atmospheres is assessed, concluding that its action may be weak in comparison with the role of corpuscular radiation.

121. K. Wurm
STRUCTURE AND DEVELOPMENT OF COMETARY TAIL
Hamburger Sternwarte, Hamburg-Bergedorf, Germany, March 31, 1964
AD 602 048

A final report on contract AF61(052)-583 presents five separate papers which deal with the study of the structure and development of ion-type cometary tails.

122. K. Wurm
DIE IONISIERUNG IN DEN KOMETEN
Icarus 1, 144 (1962)

New support is given to the theory that the ionization of cometary molecules is intrinsic to the cometary atmospheres themselves. It is argued that the ionization occurs in a limited region in front of the nucleus and that the ions are expelled in narrow rays within a small cone which is directed toward the sun. Apparently, outbursts lead to the formation of a shrinking parabolic envelope around the nucleus, which is accompanied by a closing-in of the tail rays to the tail axis. The true mechanism of ionization is still obscure.

123. K. Wurm and B. Balazs
THE HEADS OF COMETS AND THE LIFETIMES OF THEIR MOLECULES
Icarus 2, 334 (1963)

From the extension and spherical shape of the CN and $\rm C_2$ heads, an average lifetime of these molecules until ionization of the order 10^{6} seconds for unit heliocentric distance is derived. This long lifetime excludes densities of protons in the "solar wind," which exceed np = $10~\rm cm^{-3}$. (Author)

124. K. Wurm

BRIGHTNESS AND STRUCTURE OF THE NA-EMISSION IN COMETS

Hamburger Sternwarte, Hamburg-Bergedorf, Germany, March 31, 1964

AD 60h 0h8

The structure of a Na-coma is investigated. In consequence of a relatively high repulsive force originating from light pressure and a relatively high average lifetime of the atoms, a typical parabolic shape to the coma is expected.

125. A. Mammano and K. Wurm
THE AXES OF THE TYPE-I TAILS OF COMETS
Hamburger Sternwarte, Hamburg-Bergedorf, Germany, March 31, 1964
AD 604 048

The axes of the Type-I tails of comets are defined by the directions (position angles at the nucleus) to which the tail rays are "closing in."

The axes are always lagging behind the prolonged radius vector. It is shown that the directions of the tail axes are determined by the forces which regulate the outflow of the tail material from its source in the head. A direct influence of the "solar wind" on the position of the tail axes is not apparent. (Author)

126. K. Wurm

AN EXPANDING HALO OF COMET ALCOCK (1963b)
Hamburger Sternwarte, Hamburg-Bergedorf, Germany, March 31, 1964
AD 60h 0h8

On May 27/28, 1963 Comet Alcock showed an increase in light of about three magnitudes accompanied by an outbreak of gas and dust. The gas halo expanded with a velocity of one kilometer per second. The expansion of the dust material was much smaller amounting only to 50 to 100 meters per second. (Author)

127. M. J. S. Belton, J. C. Brandt and P. W. Hodge
ON THE CHARACTERISTICS OF DISTANT COMETS
Space Science Laboratory, University of California, Berkeley, California
AD 299 839

Donn's article on the characteristics of distant comets is criticized. It is shown that his arguments concerning the composition of the tails and their curvature and orientation are mutually contradictory. Brandt's use of these tails to probe the interplanetary medium is strengthened and his arguments for a transition region between the solar wind and the solar breeze at 2 AU are reinforced.

128. J. C. Brandt

ON THE STUDY OF COMET TAILS AND MODELS OF THE INTERPLANETARY MEDIUM Astrophysical Journal 133, 1091 (1961)

The theories of Bierman, Parker, and Chamberlain are applied to Osterbrock's observations of Comets Baade 1954h and Haro-Chavira 1954k.

129. J. C. Brandt

INTERPLANETARY MEDIUM PHYSICS AND COMET TAILS California University, Berkeley, 1964 AD 603 780

The programs, results, and reports resulting from research on interplanetary medium physics and comet tails are summarized.

130. S. Grundzinska
THE ABUNDANCE OF CO+ IONS IN THE TAILS OF COMETS
Ann. d'Astrophy. 23, 797 (1960)

The abundance of ${\rm CO}^+$ in the tail of Comet Bester (1947k) is found to be 30 to 40 times greater than that of CN in the head. Assuming that the abundance for CN is taken equal to that suggested for ${\rm C}_2$ by Wurm, the abundance of ${\rm CO}^+$ ions in the tail around the region of the head is estimated to be on the order of ${\rm 10}^6$ per cm³.

131. S. L. Miller
THE OCCURRENCE OF GAS HYDRATES IN THE SOLAR SYSTEM
Proc. Nat. Acad. Sci. 47, 1798 (1961)

The occurrence and properties of the hydrates in planetary, interplanetary, and cometary environment is discussed. A mixed hydrate of methane, carbon dioxide, ethane, etc. is postulated to be a likely constituent of comets.

132. K. Watson, B. C. Murray and H. Brown
THE STABILITY OF VOLATILES IN THE SOLAR SYSTEM
Icarus 1, 317 (1963)

A study of the stability of the ices of water, ammonia, carbon dioxide, and methane is extended to the rings of Saturn, the surfaces of the smaller satellites of Jupiter and Saturn, and bodies in orbits similar to those of short-period comets.

133. D. E. Blackwell and R. V. Willstrop
A STUDY OF THE MONOCHROMATIC POLARIZATION OF COMET AREND-ROLAND (1956h)
Monthly Notices of the Royal Astronomical Society 117, 590 (1957)

Polarization measurements obtained with glass and gelatin filters are reported to have isolated the continuum near 4530Å and the (0,0) CN emission band near 3850Å. The polarization curve of the continuum was found to agree reasonably with that of a metallic meteorite, and the amount of polarization in the CN measurements was consistent with a fluorescence mechanism.

134. J. D. Kraus
CBSERVATIONS AT A WAVELENGTH OF 11 METERS DURING THE CLOSE APPROACH OF
COMET AREND-ROLAND
Astronomical Journal 53, 55 (1958)

The observation of a radio source which may have been related to Comet Arend-Roland is reported.

135. J. W. Shaw
NATURAL ENVIRONMENT OF INTERPLANETARY SPACE
The Ohio State University Research Foundation, Columbus, Ohio, 1960
AD 250 230

Chapter IV of this report is devoted to solid particles in interplanetary space and includes a short description of the form and composition of comets, their orbits, and the number of observed and unobserved comets.

LABORATORY STUDIES

136. G. Herzberg, A. Monfils and B. Rosen
MOLECULAR SPECTRA IN THE REMOTE ULTRAVIOLET. INTRODUCTORY REPORT
Mem. Soc. Roy. Sci. Liege 4, 146 (1961)
OTS/SLA 63-10245

Recent progress made in vacuum ultraviolet spectroscopy is reviewed as it applies to astrophysics and theoretical molecular spectroscopy. Individual molecular spectra of the gaseous state and recent improvements in vacuum spectrographs are discussed.

137. G. Herzberg
RECENT LABORATORY INVESTIGATIONS OF MOLECULES OF ASTRONOMICAL INTEREST
Mem. Soc. Roy. Sci. Liege 28, 397 (1956)
AD 132 241

Laboratory investigations of molecules of importance to stellar atmospheres, planetary atmospheres, and comets are summarized. The discussion includes the SiC₂ molecule in carbon stars, possible detection of polyatomic molecules in stellar atmospheres, the dissociation of diatomic molecules, methods for band spectroscopy to determine temperature, the mechanism of OH production in the upper atmosphere, molecular hydrogen in planetary atmospheres, and the formation of molecules in comets.

138. G. Herzberg
LABORATORY INVESTIGATIONS OF THE SPECTRA OF INTERSTELLAR AND COMETARY
MOLECULES
Mem. Soc. Roy. Sci. Liege, Ser. 4, 15, 291 (1955)

The identification and formation of diatomic molecules in interstellar, planetary, and cometary medium is summarized. Laboratory data are presented for ${\rm CN}^+$ and ${\rm NH}^+$ and for the triatomic radicals ${\rm C}_3$ and ${\rm NH}_2$ of comets. The possible presence of HCO in interstellar medium and planetary atmospheres is discussed.

139. J. Oro
COMETS AND THE FORMATION OF BIOCHEMICAL COMPOUNDS ON THE PRIMITIVE EARTH
Nature 190, 389 (1961)

It is suggested that one important consequence of the interactions of comets with the earth is the accumulation of relatively large amounts of carbon compounds which are known to be transformed spontaneously into amino acids, purines, and other biochemical compounds. Laboratory experiments are described in verification of this theory.

140. A. E. Potter and B. Del Duca LIFETIME IN SPACE OF POSSIBLE PARENT MOLECULES OF COMETARY RADICALS Icarus 3, 103 (1964)

The space lifetimes of some possible parent molecules of radicals are estimated from laboratory measurements of absorption cross sections and the solar ultraviolet flux. The parent molecules include ethylene, acetylene, methane, cyanogen, hydrazine, ammonia, water, etc.

741. P. Wilkinson
DIATOMIC MOLECULES OF ASTROPHYSICAL INTEREST: IONIZATION POTENTIALS AND DISSOCIATION ENERGIES
Astrophysical Journal 138, 778 (1963)

A critical analysis is made of the ionization potentials and dissociation energies of 148 diatomic molecules and molecular ions of astrophysical importance. The data are summarized in a table and includes values and references for $\mathrm{C_2}$, CH , CO^+ , NH , $\mathrm{N_2}^+$, OH^+ , $\mathrm{O_2}^+$, as well as others.

142. G. J. Pontrelli
THE FORMATION AND PROPERTIES OF CERTAIN COMETARY SPECIES
Final Report on NASA Contract NASW 707, E. I. DuPont de Nemours and Company,
Wilmington, Delaware, January 1965

 C_2 , C_3 , and CH radical-radical reactions and properties observed in the continuous and flash photolysis of diacetylene are related to the postulated mechanisms of radical formation at or near the nucleus of comets.

143. G. Herzberg
THE SPECTRA AND STRUCTURES OF FREE METHYL AND FREE METHYLENE RADICALS
Proc. Roy. Soc. A 262, 291 (1961)

Vacuum ultraviolet photolysis of dimethyl mercury and diazomethane is reported to produce the spectra of the two astrophysically important radicals, CH_3 and CH_2 . Analysis of the spectra results in the determination of the electronic and geometric structure of the radicals, which can be applied to future analysis of cometary spectra.

144. G. Herzberg and J. Shoosmith

ABSORPTION SPECTRUM OF FREE CH₃ AND CD₃ RADICALS
Can. J. Phys. <u>34</u>, 523 (1956)

A preliminary study presents the absorption spectrum of CH3 as produced by the vacuum ultraviolet flash photolysis of dimethyl mercury.

145. G. Herzberg and P. A. Warsop SPECTRUM AND STRUCTURE OF THE FREE HNCN RADICAL Can. J. Phys. 41, 286 (1963)

The spectra and structure of HNCN produced in the flash photolysis of diazomethane are investigated. The 3440A band is assigned to this free radical, and its ground and first-excited state are determined.

146. G. Herzberg
LABORATORY PRODUCTION OF THE \$\times 4050 GROUP OCCURRING IN COMETARY SPECTRA;
FURTHER EVIDENCE FOR THE PRESENCE OF CH2 MOLECULES IN COMETS
Astrophysical Journal 96, 314 (1942)

The λ 4050 system occurring in cometary spectra is assigned to CH₂ (which is later refuted by Herzberg and others and assigned to C₃).

147. G. Herzberg and J. Shoosmith
SPECTRUM AND STRUCTURE OF THE FREE METHYLENE RADICAL
Nature 183, 1801 (1959)

An explanation for the previous misidentification of the λ 4050 as due to CH₂ is presented through more recent identification of the vacuum ultraviolet spectrum (λ 1400) of CH₂ produced by the photolysis of diazomethane.

148. G. W. Robinson and M. McCarty
THE PRODUCTION AND SUBSEQUENT PHOTOLYSIS OF TRANSIENT PRODUCTS FROM THE PHOTODECCMPOSITION OF DIAZOMETHANE AT 4.2°K
J. Am. Chem. Soc. 82, 1859 (1960)

Experiments on the photolysis of solid diazomethane and its isotopic modifications present evidence for the assignment of part of the observed spectra to an electronic transition of methylene in the near ultraviolet. Other products of the photolysis were trapped, and their nature is described.

149. N. Sreedhara Murthy A LABORATORY STUDY OF THE 'ROTATIONAL TEMPERATURE' OF CH⁺ ($^{1}\pi$ - $^{1}\Sigma$) COMETARY BAND Proc. Phys. Soc. 81, 1141 (1963)

The determination of the intensities of five rotational lines of the (0,0) band of the CH⁺ ($^{1}\pi$ - $^{1}\Sigma$) cometary system is reported as the result of an analysis of spectra produced through excitation by a high-frequency electrodeless discharge.

150. A. E. Douglas
LABORATORY STUDIES OF THE \$\text{\chi}\$ 4050 GROUP OF COMETARY SPECTRA
Astrophysical Journal 114, 466 (1951)

Study of the spectra obtained from laboratory production of the $\lambda\,4050$ group previously observed in cometary spectra indicates that it is due to the $^{\rm C}_3$ molecule. Isotope effects and the fine structure show that hydrogen is not present and therefore, the emission can not be due to CH $_2$ as postulated by Herzberg.

151. R. G. Bennett and F. W. Dalby
EXPTRIMENTAL OSCILLATOR STRENGTHS OF CH AND NH
J. Chem. Phys. 32, 1716 (1960)

The oscillator strengths of the A $^2\Delta$ and B $^2\Sigma$ states of CH and the A $^2\pi$ state of NH are calculated from the determination of the radiative lifetimes of the emitting species as produced by electron bombardment of methane and ammonia.

152. K. Clusius and A. E. Douglas THE λ 4050 BANDS OF THE C₃¹³ MOLECULE Can. J. Phys. 32, 319 (1954)

Excitation of the λ 4050 cometary bands with C¹³ provide very good evidence that the C₃ molecule is responsible for this emission in comets.

153. G. V. Marr and R. W. Nicholls
THE EMISSION OF THE 'A 4050A' BANDS AND AN ASSOCIATED VIOLET CONTINUUM IN
OXYACETYLENE FLAMES
Can. J. Phys. 33, 394 (1955)

A strong violet continuum is reported to appear with the $^{\rm C}_3$ bands excited in the luminous region of a fuel-rich oxyacetylene flame. Preliminary photographic intensity measurements on the spectra of the different regions of the flame show strong correlation between the $^{\rm C}_3$ bands and the continuum.

154. G. V. Marr
CONBUSTION STUDIES OF ASTROPHYSICAL SIGNIFICANCE. III.
The University of Western Ontario, London, Ontario, Canada, March 1, 1957
AD 152 611

Spectroscopic investigation of the luminous mantle of fuel-rich oxyacetylene flames encompasses band-head intensity measurements for the radicals C_2 , CH, C_3 , and CN; rotational plots for C_2 , CH, and C_3 and C_4 ; and C_5 are considered in relation. The experimental observations of C_3 are considered in relation to previous laboratory and astrophysical investigations. The formation of C_3 in comets is discussed, and the fluorescence mechanism by solar radiation is shown to account for the cometary band emission occurring in the absence of continuum emission.

155. P. S. Skell and L. D. Wescott
CHEMICAL PROPERTIES OF C₃, A DICARBENE
J. Am. Chem. Soc. 85, 1023 (1963)

The chemical reactions of carbon vapor, produced in vacuo, with olefins at liquid-nitrogen cooled surfaces are reported. The conditions of reaction have been controlled to simulate cometary ambient conditions.

156. B. Rosen

NEW INVESTIGATIONS ON THE C3 SPECTRUM AT 4050Å

Astronomical Journal 66, 383 (1961)

A brief summary of laboratory investigation of the λ 4050 system for C₃ and its importance to cometary physics is presented.

157. G. V. Marr
COMBUSTION STUDIES OF ASTROPHYSICAL SIGNIFICANCE. 1. A SURVEY OF THE
GENERAL MECHANICS OF FLAME PROPAGATION, AND OF RELEVANT EXPERIMENTAL
TECHNIQUES
The University of Western Ontario, London, Ontario, Canada, May 1, 1956
AD 101 687

The general mechanics and experimental techniques for flame propagation are presented and directed toward an effort to provide information on excitation conditions in astrophysical sources. Hydrocarbon/oxygen premixed flames are considered and it is suggested that C_2H_2/O_2 flame be used as a source of the C_3 radical which occurs in cool stars and comets.

158. G. Herzberg and D. A. Ramsay
ABSORPTION SPECTRUM OF FREE NH₂ RADICALS
J. Chem. Phys. 20, 347 (1952)

The α -bands observed in comets is assigned to NH $_2$ through preliminary investigations of the absorption spectra produced in the flash photolysis of NH $_3 \bullet$

159. G. Herzberg and D. A. Ramsay
THE ABSORPTION SPECTRUM OF FREE NH₂ RADICALS
Discussions of the Faraday Society 14, 11 (1953)

A thorough laboratory investigation of the absorption spectra of NH $_2$ as produced from the flash photolysis of NH $_3$, is presented. The structure of the spectrum and isotope shifts obtained with deuterated NH $_3$ confirm that NH $_2$ is responsible for the α -bands observed in cometary spectra.

160. D. A. Ramsay

THE ARSORPTION SPECTRA OF FREE NH AND NH $_2$ RADICALS PRODUCED BY THE FLASH PHOTOLYSIS OF HYDRAZINE

J. Phys. Chem. 57, 415 (1953)

Some preliminary results from the flash photolysis of hydrazine include the determination of the critical flash intensity above which thermal explosion occurs and the observation of the absorption spectra of the NH and NH $_2$ products of the explosion reaction.

161. D. A. Ramsay

ELECTRONIC SPECTRUM AND STRUCTURE OF THE FREE NH₂ RADICAL J. Chem. Phys. <u>25</u>, 188 (1956)

Rotational and vibrational analysis of the $\rm NH_2$ bands, produced by the flash photolysis of ammonia and hydrazine, is reported to verify the assignment of the \propto -bands of ammonia to the $\rm NH_2$ radical.

162. M. W. Feast

AN EMISSION BAND SYSTEM ATTRIBUTED TO THE MOLECULE NH+ Astrophysical Journal 114, 344 (1951)

Emission bands produced from the excitation of flowing ammonia in a hollow-cathode discharge tube are identified as NH⁺. The postulated importance of this molecule is based on previous observations of NH in cometary spectra and the dual presence of both CH and CH⁺ in cometary spectra.

163. J. A. Glasel

STABILIZATION OF MH IN HYDROCARRON MATRICES AND ITS RELATION TO COMETARY PURNOMENA

Proc. Nat. Acad. Sci. 47, 174 (1961)

The ultraviolet spectra of NH resulting from RF discharge through the low-temperature solids $\text{CH}_{\chi_4}-\text{N}_2-\text{A}$, $\text{CH}_{\chi_4}-\text{N}_2$, and NH_3-A is shown in an attempt to substantiate that free radicals such as NH can be trapped in the frozen solids which exist in comets.

164. F. O. Rice and M. Freamo

THE FORMATION OF THE IMINE RADICAL IN THE ELECTRICAL DISCHARGE J. Am. Chem. Soc. 75, 548 (1953)

The formation of the NH radical from electrical discharge in hydrazoic acid is described. Allied experiments with electrical discharge in ammonia, hydrazine, cyanic acid, and nitrogen-hydrogen mixtures failed to produce the radical.

165. L. J. Stief and V. J. DeCarlo ORIGIN OF NH(A³ π) NH(X³ Σ -) EMISSION IN COMETS Nature 205, 889 (1965)

The photolysis of hydrazine at the krypton resonance lines and at low pressures is reported to produce the λ 3360 emission spectra of NH. The spectra are that of the triplet system which has been previously observed in cometary spectra. From this, hydrazine-type compounds are postulated to be responsible for the NH emission in comets.

166. H. Schuler and A. Michel
ON NEW EMISSION BANDS FROM ELECTRIC EXCITATION OF AMMONIA
Z. Naturforsch. 10a, 1 (1955)

Two band systems appearing between 5200 and 8500Å are reported to occur upon electrical excitation of NH3 at one atmosphere.

167. W. Groth and K. H. Welge
BLITZLICHT PHOTOLYSE UND PHOTOCHEMISCHE ZERSETZUNG DES AMMONIAKS
Bull. Soc. Chim. Belg. 71, 705 (1962)

The ultraviolet flash photolysis of NH_3 to produce the NH and NH_2 radicals is described.

168. K. D. Bayes, K. H. Becker and K. H. Welge FLASH PHOTOLYSIS OF NH3 IN THE VACUUM ULTRAVIOLET Z. Naturforsch. 17a, 676 (1962)

The production of $\rm NH_2$ and $\rm NH$ radicals from the vacuum ultraviolet photolysis of gas-phase $\rm NH_3$ at wavelengths between 2200 and 1650A is reported.

169. K. H. Becker, F. Stuhl and K. H. Welge
FLASH PHOTOLYSIS AND FLUORESCENCE OF NH3 IN THE VACUUM ULTRAVIOLET
Bonn University, Germany, 1962
AD 438 411

The production, kinetics, and electronic states of the NH₂ and NH radical in the vacuum ultraviolet flash photolysis of NH₃ is described. In fluorescence studies of the primary product NH, the cl π al Σ transition, is observed, but not the 3π 3 Σ - transition.

170. H. J. Rommel
PHOTOLYSIS OF AMMONIA IN THE FAR ULTRAVIOLET
Bonn University, Germany, 1963
AD 438 494

The influence of the NH radical on the overall mechanism of the ammonia photolysis is assessed.

171. F. Stuhl and K. H. Welge
FLASH PHOTOLYSIS OF NH3 IN THE VACUUM ULTRAVIOLET. II.
Z. Naturforsch. 18a, 900 (1963)

An analysis is made of the formation of NH in the vacuum ultraviolet photolysis of NH $_3$. It is reported that NH radicals appear directly in the $^3\Sigma$ - state.

172. W. E. Groth
PHOTOCHEMICAL INVESTIGATIONS IN THE FAR ULTRAVIOLET
Bonn University, Germany, 1963
AD 416 366

The report summarizes present progress made in a series of investigations on the far ultraviolet photochemical behavior of NH3.

173. W. Groth, H. Okabe and H. J. Rommel PRIMARY PROCESSES OF NH3 PHOTOLYSIS AT 1470A Z. Naturforsch. 19a, 507 (1964)

The photolysis of NH $_3$ at 1470Å is investigated and found to proceed to 12 percent by reaction to NH radicals and to 38 percent by reaction to NH $_2$ radicals.

174. K. H. Becker and K. H. Welge THE INVESTIGATION OF THE FLUORESCENCE AND PHOTOCHEMICAL PRIMARY PROCESSES IN THE VACUUM UV BY NH_3 , N_2H_4 , PH_3 AND THE REACTIONS OF THE ELECTRONICALLY EXCITED RADICALS NH^* ($^1\pi$), NH^* ($^3\pi$), PH^* ($^3\pi$) Z. Naturforsch. 19a, 1006 (1964)

(In German)

175. E. Fink and K. H. Welge LIFETIME OF THE ELECTRONIC STATES $N_2(c^3\pi u)$, $N_2^+(B^2\Sigma^+u)$, $N_1(c^2\pi)$, $N_2^+(B^2\Sigma^+u)$, $N_2^+(B^2\Sigma^+u)$, $N_1(c^2\pi)$, $N_2^+(B^2\Sigma^+u)$

(In German)

176. K. D. Beyer and K. H. Welge
PHOTODISSOCIATION OF H₂, N₂, O₂, NO, CO, H₂O, CO₂ AND NH₃ IN EXTREME VACUUM
ULTRAVIOLET
Z. Naturforsch 19a, 19-28 (1964)

(In German)

177. K. H. Becker and K. H. Welge
FLUORESCENCE OF NH RADICALS DURING THE PHOTODISSOCIATION OF NH3 IN THE
VACUUM ULTRAVIOLET
Z. Naturforsch 18a, 600-3 (1963)

(In German)

178. A. E. Douglas and P. M. Routly
THE SPECTRUM OF THE CN⁺ MOLECULE
Astrophysical Journal 119, 303 (1954)

It is reported that two band systems due to two $^1\Sigma$ - $^1\Sigma$ rotational transitions of the astrophysically-important CN+ molecule have been produced from a discharge through helium mixed with a trace of C₂N₂. Four other unidentified bands are also reported.

179. R. G. Bennett and F. W. Dalby
EXPERIMENTAL OSCILLATOR STRENGTH OF THE VIOLET SYSTEM OF CN
J. Chem. Phys. 36, 399 (1962)

The electron bombardment of HCN to produce the emission spectrum of CN is described. The oscillator strength of the $\rm B^2\Sigma$ state is determined from a direct measurement of the spontaneous radiative lifetime.

180. D. E. Paul and F. W. Dalby KINETICS OF DISAPPEARANCE OF THE CN RADICAL FORMED FROM C_2N_2 J. Chem. Phys. 37, 592 (1962)

The flash photolysis of C_2N_2 to produce CN radicals is utilized in a kinetic study of the rate of disappearance of CN radical in cyanogen, cyanogen chloride, hydrogen, and hydrocarbons. The emission spectrum of the CN radical is found to arise from a fluorescence of the CN radical and not from formation of radical in its excited state.

181. A. E. Douglas and P. M. Routley
ON THE SPECTRA OF THE CN⁺ AND CN MOLECULES
Astrophysical Journal 117, 461 (1953)

The spectra of CN⁺ and CN observed from C_2N_2 in a helium discharge tube are reported. The band origins and rotational constants of $^1\Sigma$ - $^1\Sigma$ transition of CN⁺ are listed and two new band systems of neutral CN, occurring in the same region as the CN⁺, are identified.

182. R. L. Brown and H. P. Broida
SPECTRAL STUDY OF ACTIVE NITROGEN FLAMES EXHIBITING CN "TAIL" BANDS
J. Chem. Phys. 41, 2053 (1964)

Flames of active nitrogen and simple chlorinated hydrocarbons are reported to exhibit several zones which can be defined by their differences in intensity and spectra. One zone has been found to exhibit that emission which arises from the $^{\rm B2}\Sigma$ state of CN. The mechanism of its occurrence is postulated.

183. R. N. Dixon and R. W. Nicholls
AN EXPERIMENTAL STUDY OF THE BAND INTENSITIES IN THE CN RED SYSTEM
Can. J. Phys. 36, 127 (1958)

Experimental band intensities are reported for the CN red system obtained from an active nitrogen - carbon tetrachloride source.

184. E. A. Ballik and D. A. Ramsay GROUND STATE OF THE C₂ MOLECULE J. Chem. Phys. 31, 1128 (1959)

The rotational and vibrational analysis of C_2 band system as observed in the near infrared spectrum of the emission from a carbon furnace indicate that the $3\sum_g$ -state is the ground state of C_2 in the gas phase.

185. E. A. Ballik and D. A. Ramsay THE A' $^3\Sigma_g$ - X' $^3\pi_u$ BAND SYSTEM OF THE C₂ MOLECULE Astrophysical Journal 137, 1 (1963)

The detail of the vibrational and rotational lines of the infrared band system of the $\rm C_2$ molecule is presented and a possible mechanism is advanced to explain the presence of the $\rm C_2$ Swan bands and absence of the $\rm C_2$ Phillips bands in cometary spectra.

186. E. A. Ballik and D. A. Ramsay

AN EXTENSION OF THE PHILLIPS SYSTEM OF C2 AND A SURVEY OF C2 STATES

Astrophysical Journal 137, 34 (1963)

Nine new bands of the Phillips system of C_2 are presented and measurements on three other bands are extended to higher J values, with new molecular constants evaluated for both states. A general summary of molecular constants for all known states of C_2 and a diagram of the potential curves are given also.

187. V. J. DeCarlo
STUDY IN COMETARY ASTROPHYSICS
NASA Contractor Report, National Aeronautics and Space Administration,
Washington, D. C., July 1964

 C_2 emission observed from the vacuum ultraviolet photolysis of low-pressure methane and acetylene is reported to coincide with the Swan bands observed in cometary spectra. A second investigation of the ultraviolet photolysis of ice (H₂O) produced unidentified and unreproducible emission bands and continuum, each occurring between $\sim \lambda$ 4000-4700.

188. L. J. Stief, V. J. DeCarlo and R. J. Mataloni
VACUUM ULTRAVIOLET PHOTOLYSIS OF ACETYLENE
Melpar, Inc., Falls Church, Va., November 1964, (submitted to J. Chem.
Phys. for publication)

In a study initiated by a program in cometary astrophysics, the photolysis of acetylene at the krypton and xenon resonance lines is described. Evidence is presented for the production of a long-lived excited state of acetylene which at low pressure leads to the formation of the excited C_2 . The spectra of the C_2 emission at the krypton and xenon lines are consistent with the Swan bands previously observed in cometary spectra. An excited-molecule mechanism for the photolytic reactions is postulated.

189. L. J. Stief and V. J. DeCarlo CRIGIN OF THE $C_2(A^3\pi_g)$ ——— $C_2(X^3\pi_u)$ EMISSION IN COMETS Melpar, Inc., Falls Church, Va., November 1964 (submitted to Nature for publication)

From the results of laboratory investigation, it is postulated that the C_2 Swan bands observed in cometary spectra are due to the photodissociation of acetylene or an acetylene-type molecule.

190. G. V. Marr
COMBUSTION STUDIES OF ASTROPHYSICAL SIGNIFICANCE. II. A SURVEY OF
CHEMICAL KINETICS FOR PREMIXED HYDROCARBON/OXYGEN FLAMES
The University of Western Ontario, London, Ontario, Canada, July 1, 1956
AD 110 208

The trend of possible reaction processes thought to be occurring in premixed hydrocarbon/oxygen flames is outlined as an aid in the study of reaction processes occurring in astrophysical sources such as comets, etc. The collision processes involved in conventional kinetics are applied to combustion processes. The relevant parameters are defined and an example of the analysis of the steady state departure from thermal equilibrium is given for an idealized flame. The available techniques for the analysis of the reactions occurring are outlined and considered in the postulation of probable reaction mechanisms which are capable of forming the excited C_2 , CH, HCO, and CH free radicals in premixed hydrocarbon flames.

191. R. G. Bennett and F. W. Dalby
EXPERIMENTAL OSCILLATOR STRENGTH OF COMET-TAIL SYSTEM OF CO⁺
J. Chem. Phys. 32, 1111 (1960)

The electron bombardment of CO is utilized to produce the emission spectra of the comet-tail system of ${\rm CO}^+$ and to measure the radiative lifetimes of its upper electronic state.

192. D. Robinson and R. W. Nicholls
INTENSITY MEASUREMENTS ON THE CO+ COMET TAIL, AND THE BO \(\alpha \) AND \(\beta \)
MOLECULAR BAND SYSTEMS
Proc. Phys. Soc. 75, 817 (1960)

The photoelectric measurement of the relative band intensities of the comet tail system ($A^2\pi - X^2\pi$) of CO⁺ and the BO α and β molecular band systems is reported. The study is a part of a program for the establishment of intensity and transition probability data for molecular band systems of astrophysical, aeronomical, and combustion interest.

193. R. W. Nicholls
TRANSITION PROBABILITIES OF MOLECULAR BAND SYSTEMS
The University of Western Ontario, Department of Physics, London, Canada,
June 1962
N63-18009

Franck-Condon factors calculated to high vibrational quantum numbers are presented for vacuum ultraviolet transitions in N_2 , N_2^+ , O_2 , CO, and CO $^+$.

194. H. I. S. Ferguson, R. W. Nicholls and L. Herman
EXCITATION OF THE FIRST NEGATIVE SYSTEM OF O2 BY PROTON BEAM IN AIR AND
OXYGEN
Colloq. Spectros. Intern. 9th, Lyons, 1961 2, 60-83 (1962)

An account is given of the excitation of the bands of the first positive system of 0_2 ⁺ by a beam of 40KeV protons. Experiments with nitrogen, acetylene, and water vapor are also reported. The experiments indicate that a direct excitation occurs with charge transfer.

195. D. Robinson and R. W. Nicholls INTENSITY MEASUREMENTS ON THE ${\rm O_2}^+$ SECOND NEGATIVE, CC ÅNGSTRÖM AND THIRD POSITIVE, AND NO χ AND β MOLECULAR BAND SYSTEMS Proc. Phys. Soc. 71, 957 (1958)

In an effort to obtain intensity measurements for band systems of astrophysical interest, the relative band intensities are reported for the second negative system of 0_2^+ , the Ångström and third positive systems of CO, and the γ and β systems of NO.

196. R. G. Turner and R. W. Nicholls AN EXPERIMENTAL STUDY OF BAND INTENSITIES IN THE FIRST POSITIVE SYSTEM OF N₂ Can. J. Phys. 32, 468 (1954)

Vibrational transition probabilities are interpreted from the integrated intensities of 52 bands of the N_2 first positive system.

197. R. G. Turner and R. W. Nicholls AN EXPERIMENTAL STUDY OF PAND INTENSITIES IN THE FIRST POSITIVE SYSTEM OF N2 Can. J. Phys. $\underline{32}$, 475 (1954)

Band intensities of the first positive system of nitrogen are used together with theoretical transition probabilities to determine the dependence of the electronic transition moment on the internuclear separation.

198. L. V. Wallace and R. W. Nicholls THE INTERPRETATION OF UNTENSITY DISTRIBUTIONS IN THE N₂ SECOND POSITIVE AND ${\rm N_2}^+$ FIRST NEGATIVE BAND SYSTEMS J. Atm. Terr. Phys. 7, 101 (1955)

The variation of the electronic transition moments is compared with the observed internuclear separations, and the effect of this variation upon the array of vibrational transition probabilities is assessed.

199. R. G. Bennett and F. W. Dalby EXPERIMENTAL DETERMINATION OF THE OSCILLATOR STRENGTH OF THE FIRST NEGATIVE BANDS OF ${\rm N_2}^+$ J. Chem. Phys. 31, 434 (1959)

Investigation of the electron bombardment of N₂ has led to the determination of the radiative lifetime of the upper state of the λ 391 μ transition of N₂⁺ and its corresponding oscillator strength. The radiative lifetime of the upper state of the λ 3371 transition of the second positive system of N₂ is found to depend on the conditions of excitation.

200. J. A. Glasel

THE BOMBARDMENT OF SOLID D₂0 BY LOW-ENERGY ELECTRONS UNDER ASTRONOMICAL CONDITIONS

Proc. Nat. Acad. Sci. 48, 491 (1962)

The rate of production of D_2 from the electron bombardment of solid $\mathrm{D}_2\mathrm{O}$ is measured with a helium mass spectrometer leak detector in an attempt to assess the effect of astronomical conditions on simple molecules in condensed forms.

201. R. Berger
THE PROTON IRRADIATION OF METHANE, AMMONIA, AND WATER AT 77°K
Proc. Nat. Acad. Sci. <u>17</u>, 1434 (1961)

The bombardment of low-temperature mixtures with 12-MeV protons to produce acetone, urea, and acetamide is described. The significance of this reaction is presented through postulated processes to explain the presence of NH $_3$ and CO $_2$ in comets.

202. R. G. Bennett and F. W. Dalby
EXPERIMENTAL DETERMINATION OF THE OSCILLATOR STRENGTH OF THE VIOLET
SYSTEM OF OH
J. Chem. Phys. 40, 1414 (1964)

The electron bombardment of CH $_3$ OH and H $_2$ O to produce the violet emission system of OH is used in the determination of the radiative lifetime of the A 2 π _____ X 2 Σ transition of OH and the corresponding oscillator strength.

PLANNED SPACE INVESTIGATION

203. P. Swings
POSSIBLE CONTRIBUTIONS OF SPACE EXPERIMENTS TO COMETARY PHYSICS
Smithsonian Institution Astrophysical Observatory, Cambridge, Mass., 1962
N63-14253

Suggested plans for space investigations of comets are presented and include discussions of proposed far ultraviolet observations, artificial comets, an orbiting artificial cometary nucleus, and a comet probe. A review of the physical characteristics and past astronomical observations of the spectra of comets is also discussed.

204. D. L. Roberts
THE SCIENTIFIC OBJECTIVES OF DEEP SPACE INVESTIGATIONS. COMETS
IIT Research Institute, 10 West 35th Street, Chicago, Ill., March 9, 1964
N64-19569

Deep space measurements which can provide useful data on comets are discussed along with a description of the comets of interest and the planned space missions.

205. H. C. Corben
REMARKS ON A COMET PROBE
Space Age Astronomy, p. 380, Academic Press, New York, N. Y., 1962

The requirements for launching a comet probe are given in a brief and preliminary assessment.

206. COMET INTERCEPT STUDY
Space Technology Laboratories, Inc., Redondo Beach, California
N63-16229

The study of the requirements and feasibility of a comet intercept mission includes a compilation of the properties of 31 short-term comets, the determination of the injection energies for each of the comets at a suitable launch period, the calculation of the distance of the comet from the earth at intercept, computation of the transit and flight times, determination of necessary guidance requirements, suggestions of possible scientific instruments, and an assessment of the capabilities of available boosters.

207. FAR INTERPLANETARY AND INTERSTELLAR PROBES Space/Aeronautics, January 1964, p. 98

The programs and probes planned for interplanetary and interstellar investigation are discussed. A Mariner-type craft is stated to be intended for cometary missions, which will probably not be possible before the late 60's or during the 70's. The primary obstacle to a cometary probe is reportedly more precise guidance equipment.

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